

## Quantum Mechanics

Quantisation  $L = I\omega = m_e v_n r_n = n\hbar$

Balmer Series  $v = \frac{m_e e^4}{8\varepsilon_0^2 h^3} \left( \frac{1}{2^2} - \frac{1}{m^2} \right)$

Heisenberg  $\Delta x \Delta p \geq \frac{\hbar}{2}$

Dispersion  $\omega = \frac{\hbar k^2}{2m}, \omega = ck$

Schrödinger  $-\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V(x,t)\psi(x,t) = i\hbar \frac{\partial \psi(x,t)}{\partial t}$

Tunnelling Alpha decay (atom potential)  
Ammonia molecule (nitrogen atom penetrates plane of H)  
Tunnel diode (variable height voltage potential barrier)  
STM (tunnelling = distance)

Quantum Rotator  $J = j(j+1)\hbar^2$

Quantum Oscillator  $E_n = (n + \frac{1}{2})\hbar\omega$

## Statistical Mechanics

Perfect Gas  $pV = NkT, p = nkT$

First Law  $dU = dQ + dW$

Constant V  $dU = dQ = (\frac{\partial Q}{\partial T})_V dT = C_V dT$

Constant P  $(\frac{\partial Q}{\partial T})_P = C_P = C_V + R$

Adiabatic  $dQ = 0, pV^{\frac{C_P}{C_V}} = C$

Boltzmann  $p(E)dE = A g(E) \exp(-\frac{E}{kT})dE$

Kinetic Theory Perfect gas (equation of state, no heat in a Joule expansion)  
Gases consist of small, hard, independent particles

Velocity randomly distributed

Temperature related to KE

1D Maxwell  $f_1(v_x)dv_x = \sqrt{\frac{m}{2\pi kT}} e^{-mv_x^2/2kT} dv_x$

3D Maxwell  $f(v)dv = (\frac{m}{2\pi kT})^{\frac{3}{2}} 4\pi v^2 e^{-mv^2/2kT} dv$

$$\frac{1}{2} mv^2 = \frac{3}{2} kT, \bar{v} = \sqrt{\frac{8kT}{\pi m}}$$

Pressure  $p = nmv_x^2 = \frac{1}{3} nm\bar{v}^2$

Partial Pressures  $\frac{\sigma_{M_2}}{\sigma_{M_1}} = \left(\frac{M_1}{M_2}\right)^{1/2}, p = p_1 + p_2$

Flux Density  $J = \frac{\Delta N}{\Delta t} = n \int_0^\infty v_x f_1(v_x) dv_x = \frac{1}{4} n \bar{v}$

Mean Free Path  $\lambda = \frac{1}{\pi a^2 n}, p(x) = \exp(-\frac{x}{\lambda}), R = \lambda \sqrt{N}$

Equipartition  $E = \frac{f}{2} kT, (\frac{\partial Q}{\partial T})_v = C_v = \frac{f}{2} R$

Second Law No process exists whose sole effect is the transfer of heat from a cooler to a hotter body  
 $\ln g = N \ln(Q/N), Q$  quanta, N boxes, g = permutations  
Net energy exchange leads to an increase in the most probable

number of ways in which the system can arrange itself

$$S = k \ln g, \frac{1}{T} = \frac{\partial S}{\partial E}$$

## Quantum Statistics

Raleigh-Jeans  
Body Radiation  
Planck Black Body Radiation

$$dN = \frac{8\pi v^2}{c^3} dv, u(v)dv = \bar{E}dN = \frac{8\pi^2 kT}{c^3} dv$$

$$u(v)dv = \frac{8\pi v^3}{c^3} \frac{1}{\exp(hv/kT)-1} dv$$

$$u = \int_0^\infty u(v)dv = \left(\frac{8\pi^5 k^4}{15 c^3 h^5}\right) T^4$$

Schottky Anomaly

A maximum in the heat capacity, where you run out of energy levels in which to store heat

Excitation Condition

Einstein Solid Heat Capacity

Debye Solid Heat Capacity

$$T \gg \frac{E}{k}$$

$$C = 3R \left(\frac{hv_E}{kT}\right)^2 \frac{\exp(hv_E/kT)}{\left(\exp(hv_E/kT)-1\right)^2}$$

$C \propto VT^3$  at low T by finding photon energies that will fit into a solid